First-principle theory-based scaling of the SOL width in limited tokamak plasmas and comparison with experiments

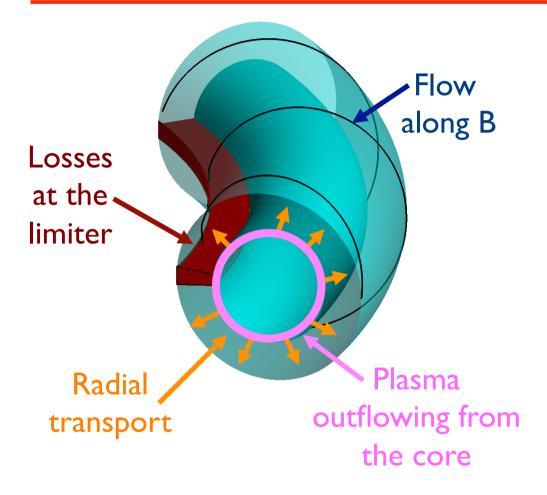
Paolo Ricci

F. Halpern, S. Jolliet, J. Loizu, and A. Mosetto,

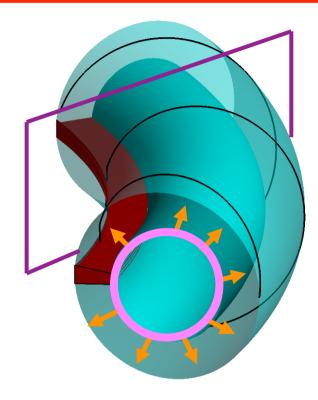
Centre de Recherches en Physique des Plasmas École Polytechnique Fédérale de Lausanne, Switzerland

How can we develop a first-principle scaling of the SOL width? The first step: simulations capturing SOL key features Interpretation of the simulation results to get the SOL width scaling How do our theoretical estimates agree with experimental data?

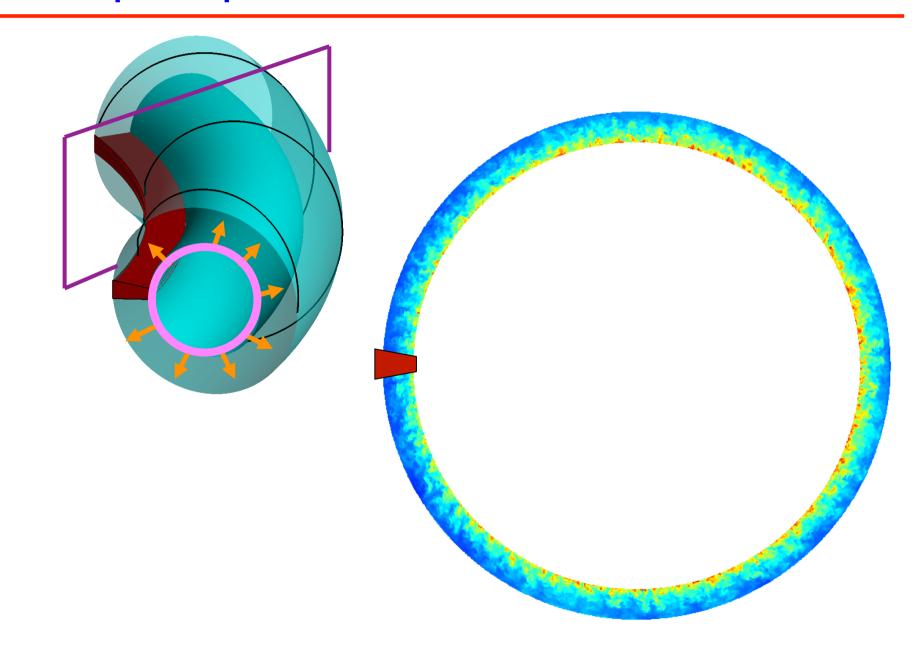
First-principle full-scale 3D SOL simulations

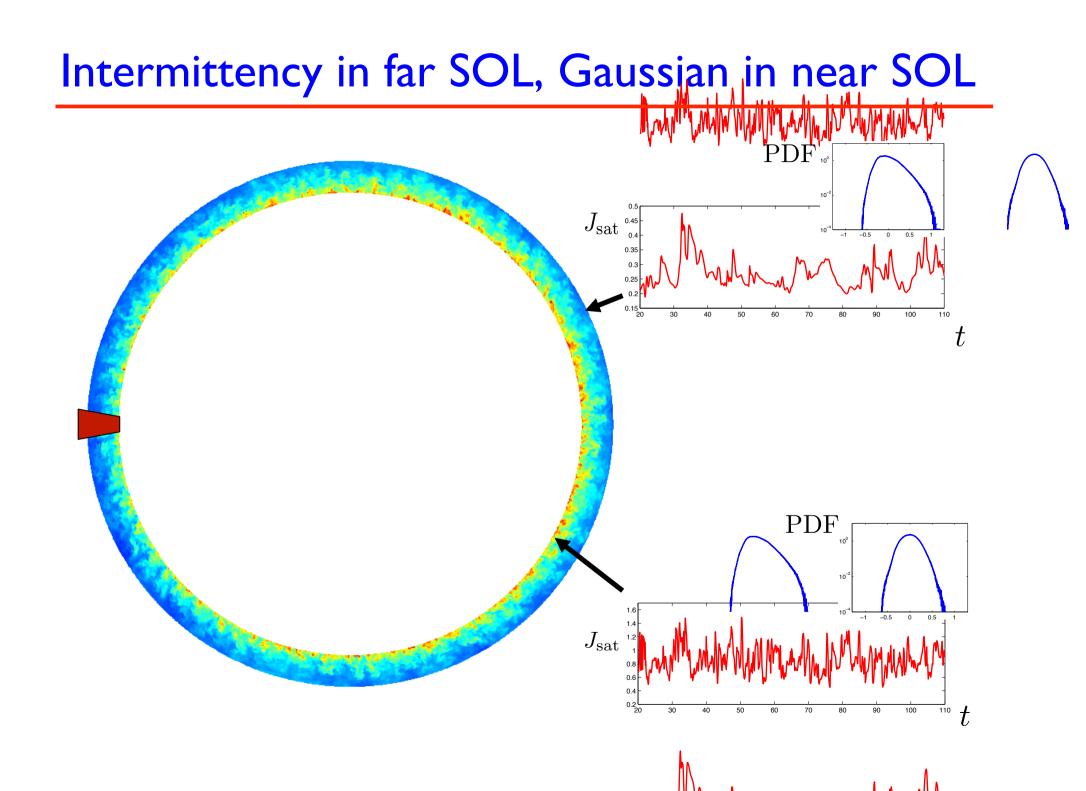


First-principle full-scale 3D SOL simulations

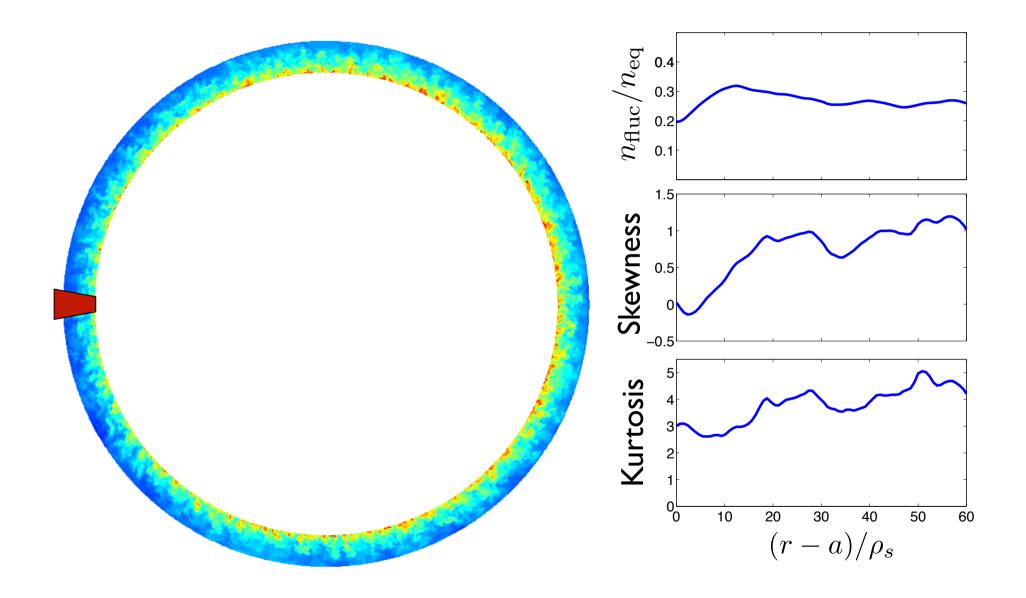


First-principle full-scale 3D SOL simulations

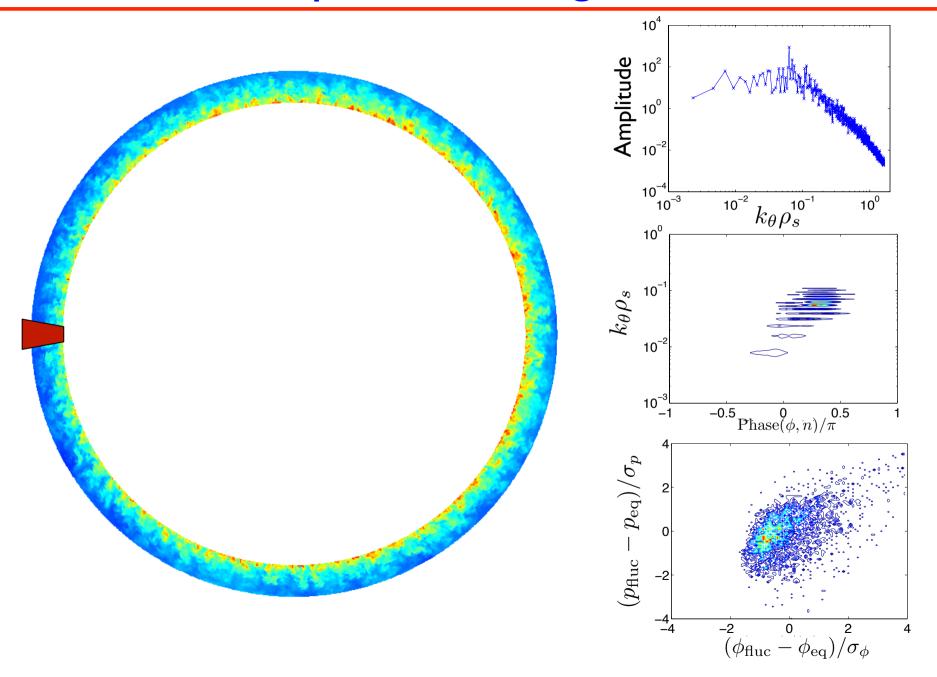




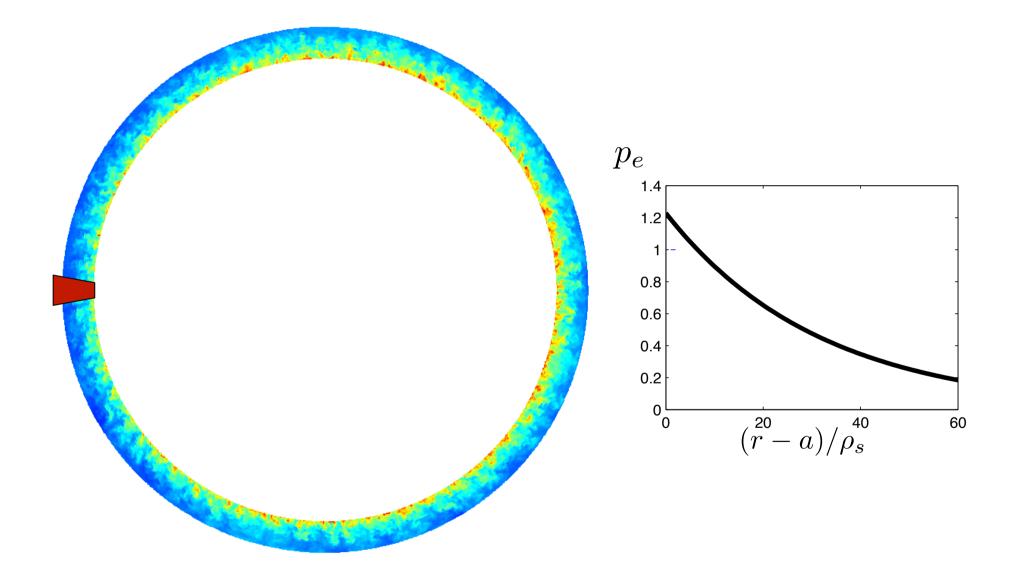
High fluctuation level, skewed PDF



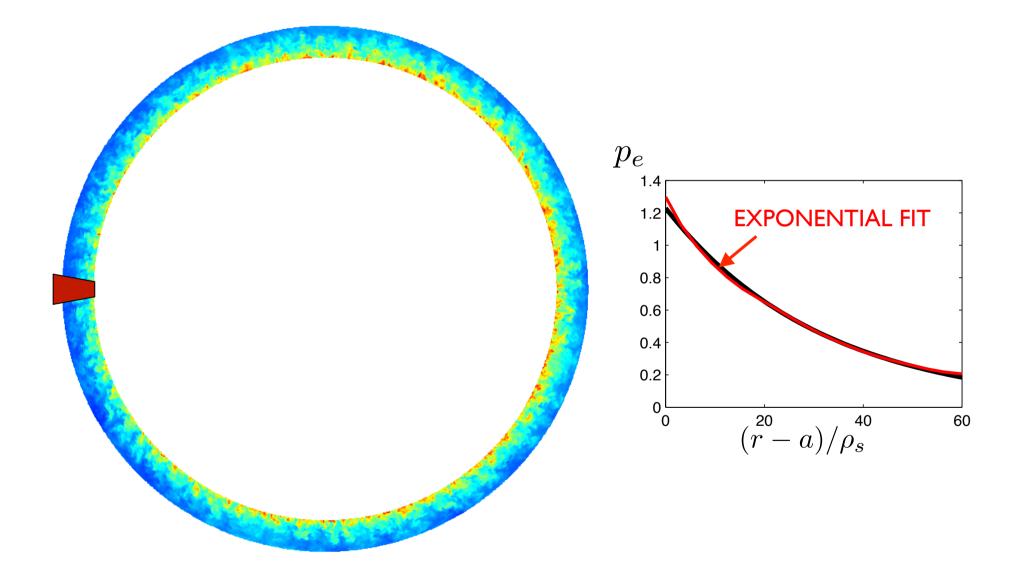
Macroscopic ballooning turbulence



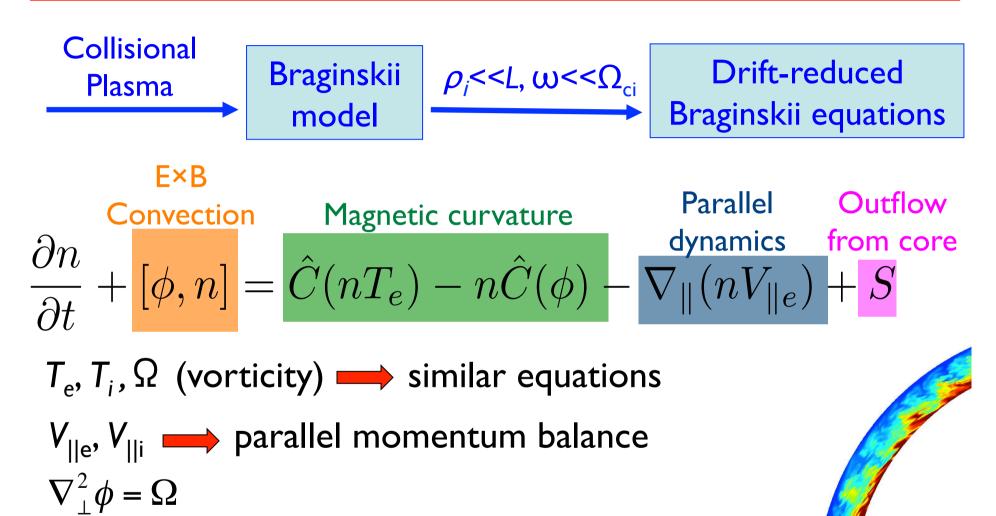
Pressure profile fitted with an exponential



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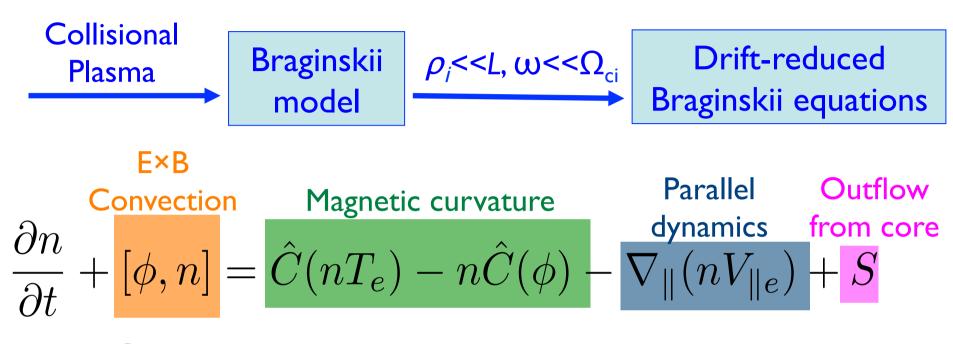


The GBS code, a tool to simulate SOL turbulence



We derived a new, first-principle, set of boundary conditions, generalizing Bohm-Chodura

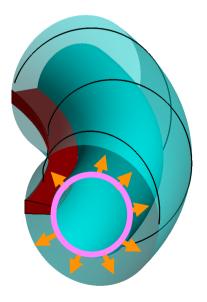
The GBS code, a tool to simulate SOL turbulence



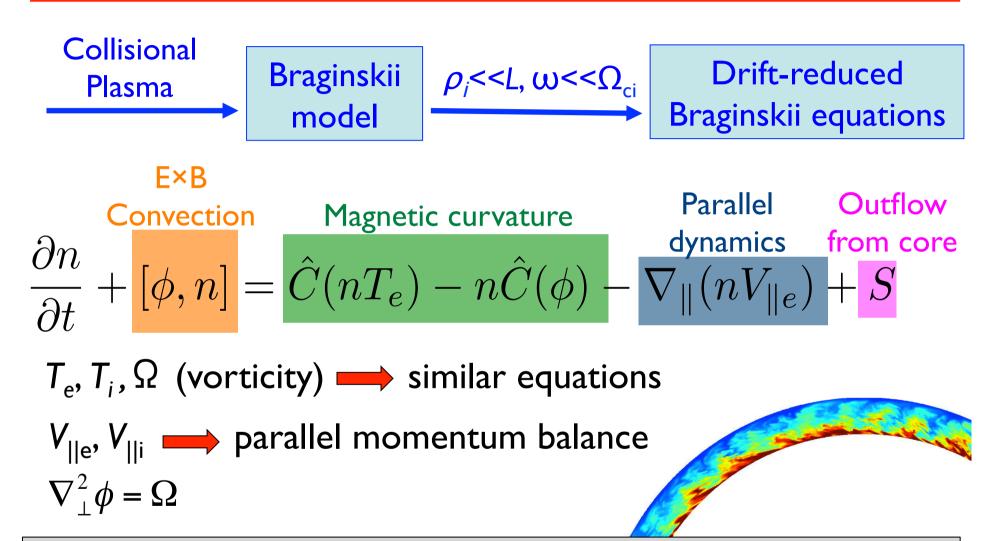
 $T_{\rm e}, T_{\rm i}, \Omega$ (vorticity) \implies similar equations

 $V_{||e}, V_{||i} \implies$ parallel momentum balance $\nabla_{\perp}^2 \phi = \Omega$

Solved in 3D, dynamics resulting from: plasma outflow, turbulent transport, and parallel losses



The GBS code, a tool to simulate SOL turbulence



Simulations contain drift physics, turbulence (ballooning modes, drift waves, ...), blobs, parallel flows, sheath losses...

The key questions

• How is the SOL width established?

• The differences between LFS and HFS limited configurations?

• What determines the SOL electrostatic potential?

• Are there mechanisms to generate toroidal rotation in the SOL?

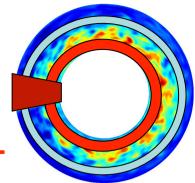
The key questions

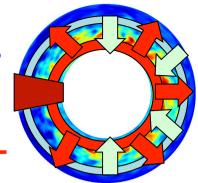
• How is the SOL width established?

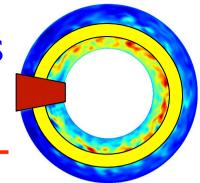
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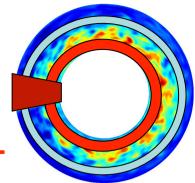
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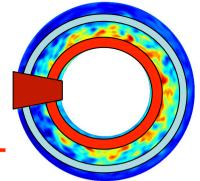
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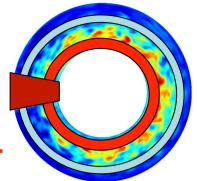






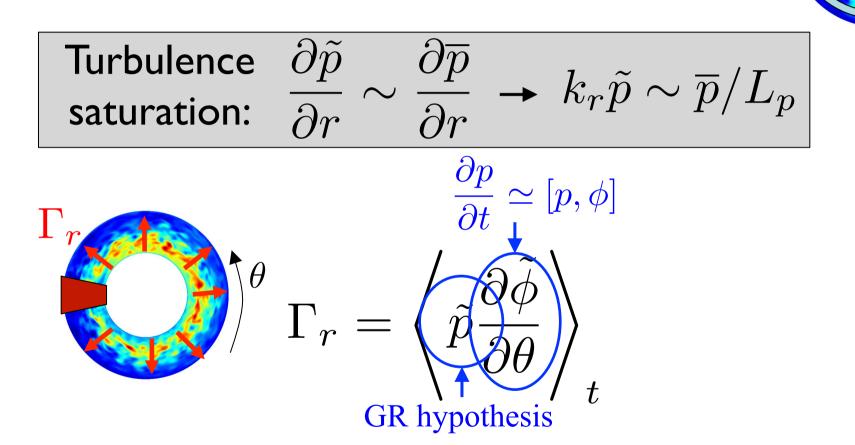


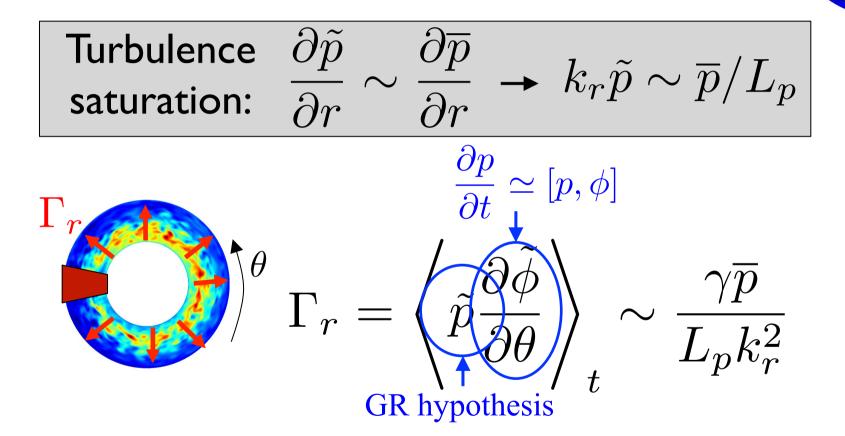
Turbulence $\frac{\partial \tilde{p}}{\partial r} \sim \frac{\partial \overline{p}}{\partial r} \rightarrow k_r \tilde{p} \sim \overline{p}/L_p$ saturation: $\frac{\partial \tilde{p}}{\partial r} \sim \frac{\partial \overline{p}}{\partial r}$



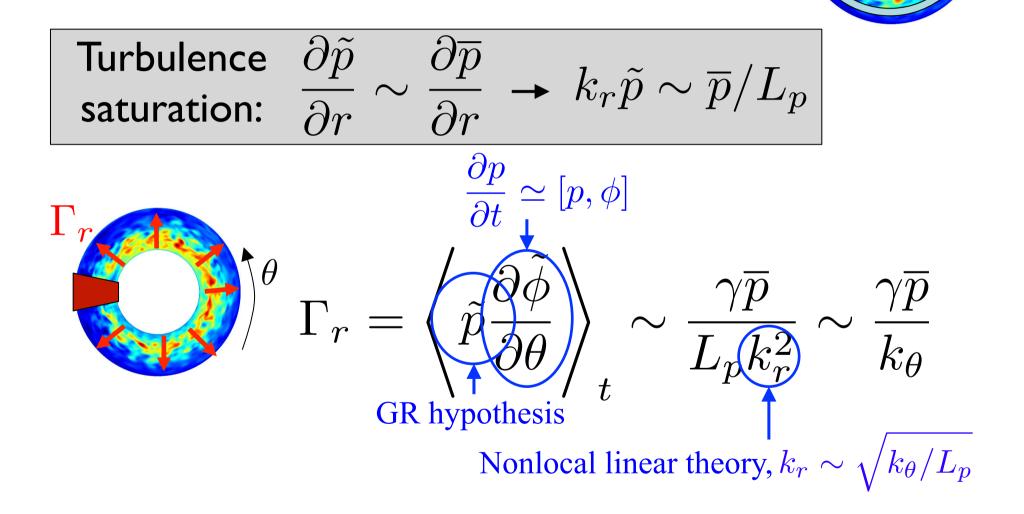
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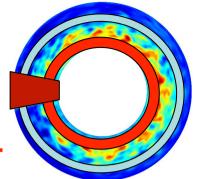
$$\Gamma_{r} \longrightarrow \theta \quad \Gamma_{r} = \left\langle \tilde{p} \frac{\partial \tilde{\phi}}{\partial \theta} \right\rangle_{t}$$





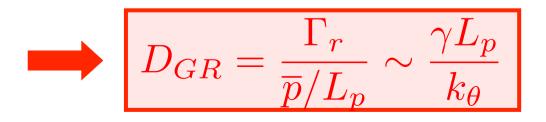
 $\frac{\partial \tilde{p}}{\partial r} \sim \frac{\partial \overline{p}}{\partial r} \rightarrow k_r \tilde{p} \sim \overline{p}/L_p$ Turbulence saturation: $\frac{\partial p}{\partial t}$ $\simeq [p, \phi]$ θ **GR** hypothesis Nonlocal linear theory, $k_r \sim \sqrt{k_{\theta}/L_p}$





Turbulence $\frac{\partial \tilde{p}}{\partial r} \sim \frac{\partial \overline{p}}{\partial r} \rightarrow k_r \tilde{p} \sim \overline{p}/L_p$ saturation: $\frac{\partial \tilde{p}}{\partial r} \sim \frac{\partial \overline{p}}{\partial r}$

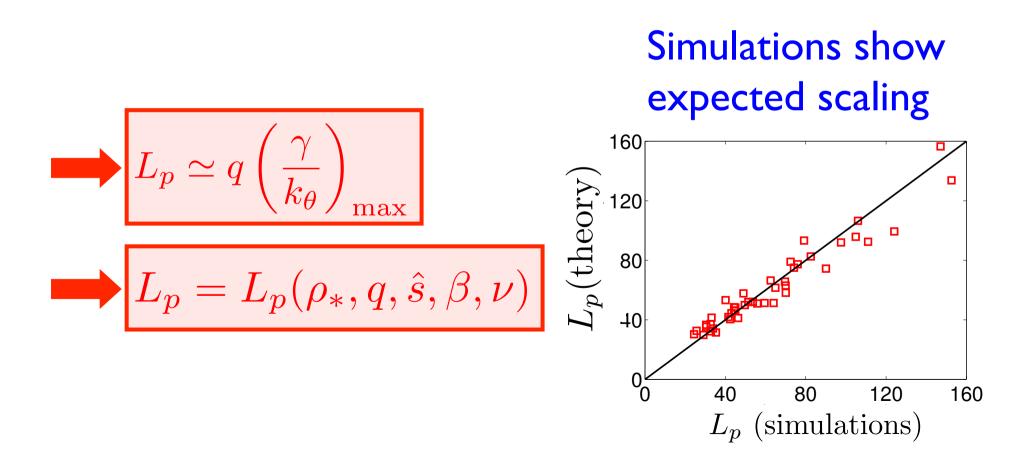
 $\int \hat{\theta} \Gamma_r = \left\langle \tilde{p} \frac{\partial \tilde{\phi}}{\partial \theta} \right\rangle_t \sim \frac{\gamma \overline{p}}{L_p k_r^2} \sim \frac{\gamma \overline{p}}{k_\theta}$



SOL width – operational parameter estimate

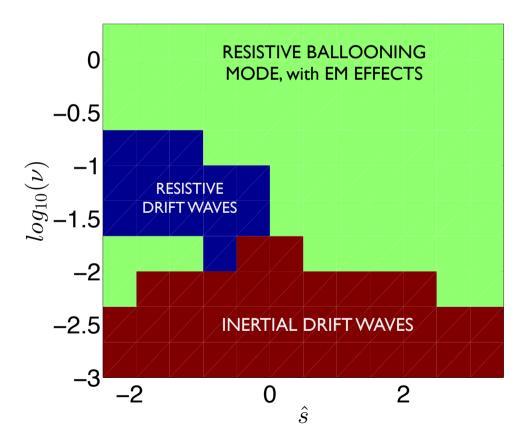
Balance of perpendicular transport and parallel losses

$$\star \frac{d\Gamma_r}{dr} \sim L_{\parallel} \underset{\text{Bohm's}}{\star} \frac{nc_s}{qR}$$



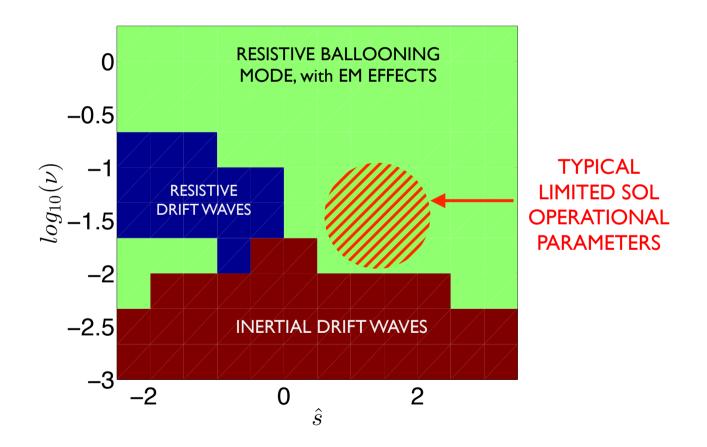
SOL turbulent regimes

Instabilities driving turbulence depends mainly on q, ν, \hat{s} .



SOL turbulent regimes

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$$L_p \simeq q \left(\frac{\gamma}{k_{\theta}}\right)_{\max}$$

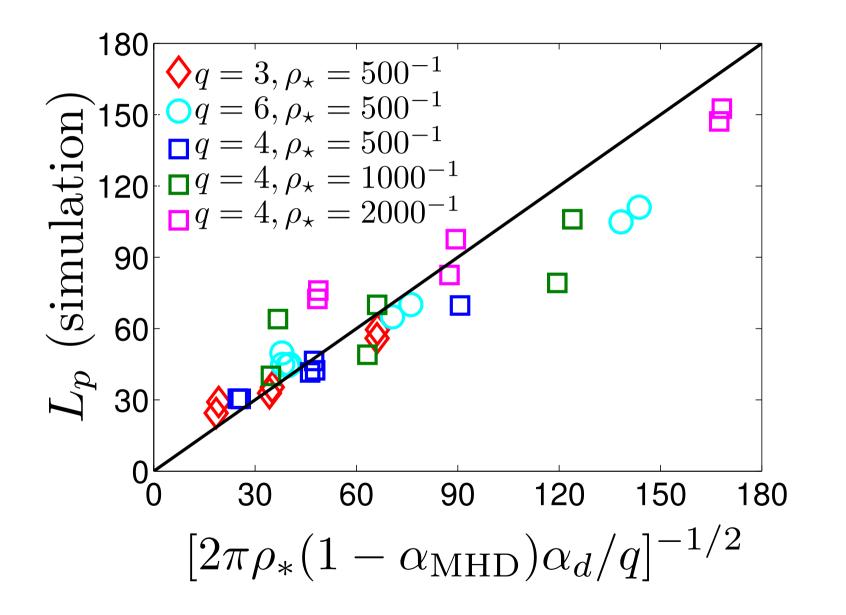
$$L_p \simeq q \left(\underbrace{\gamma}_{k\theta} \right)_{\text{BM}} \gamma \sim \gamma_b = \sqrt{2R/L_p} \left(\underbrace{\gamma}_{k\theta} \right)_{\text{BM}} k_{\theta} \sim k_b = \sqrt{\frac{1 - \alpha_{\text{MHD}}}{\nu q^2 \gamma_b}}$$

$$L_p \simeq q \left(\underbrace{\gamma}_{k\theta} \right)_{\text{BM}} \gamma \sim \gamma_b = \sqrt{2R/L_p} \\ \underbrace{k_{\theta}}_{\text{BM}} k_{\theta} \sim k_b = \sqrt{\frac{1 - \alpha_{\text{MHD}}}{\nu q^2 \gamma_b}}$$

$$L_p = [2\pi\rho_*(1 - \alpha_{\rm MHD})\alpha_d/q]^{-1/2}$$

$$L_p \simeq q \underbrace{\left(\begin{array}{c} \gamma \\ k_{\theta} \end{array}\right)}_{\text{BM}} \gamma \sim \gamma_b = \sqrt{2R/L_p} \\ \downarrow p \simeq q \underbrace{\left(\begin{array}{c} \gamma \\ k_{\theta} \end{array}\right)}_{\text{BM}} \lambda_{\theta} \sim k_b = \sqrt{\frac{1 - \alpha_{\text{MHD}}}{\nu q^2 \gamma_b}} \\ \downarrow p = [2\pi \rho_*(1 - \alpha_{\text{MHD}})\alpha_{\theta}/q]^{-1/2} \\ \downarrow q \sim q^2 \beta R/L_p \\ \downarrow q \sim q^2 \beta R/L_p \end{aligned}$$

Simulations agree with ballooning estimates



Good agreement with multi-machine measurements

The ballooning scaling, in SI units:

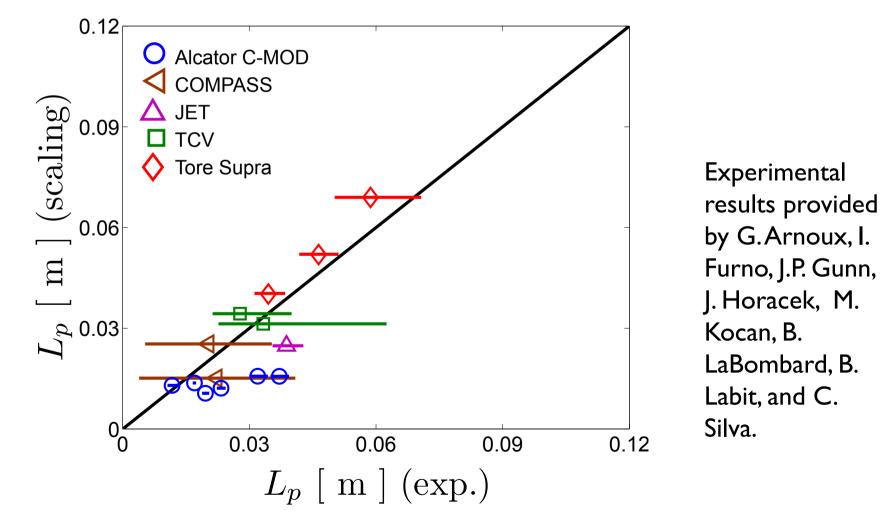
$$L_p \simeq 7.22 \times 10^{-8} q^{8/7} R^{5/7} B_{\phi}^{-4/7} T_{e,\text{LCFS}}^{-2/7} n_{e,\text{LCFS}}^{2/7} \left(1 + \frac{T_{i,\text{LCFS}}}{T_{e,\text{LCFS}}} \right)^{1/7}$$

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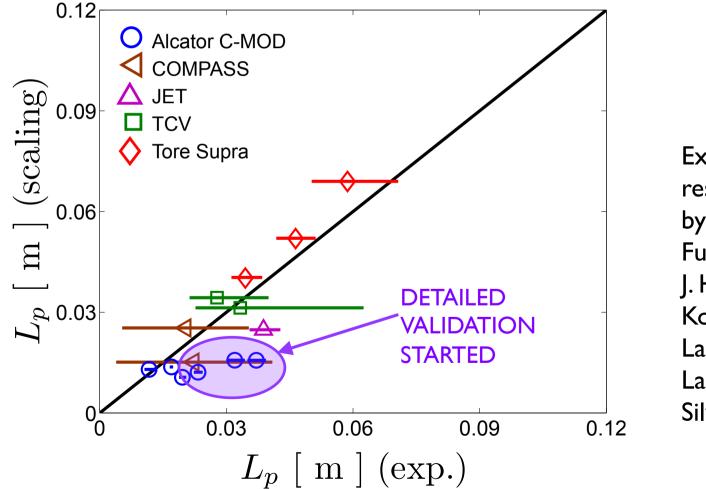
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Good agreement with multi-machine measurements

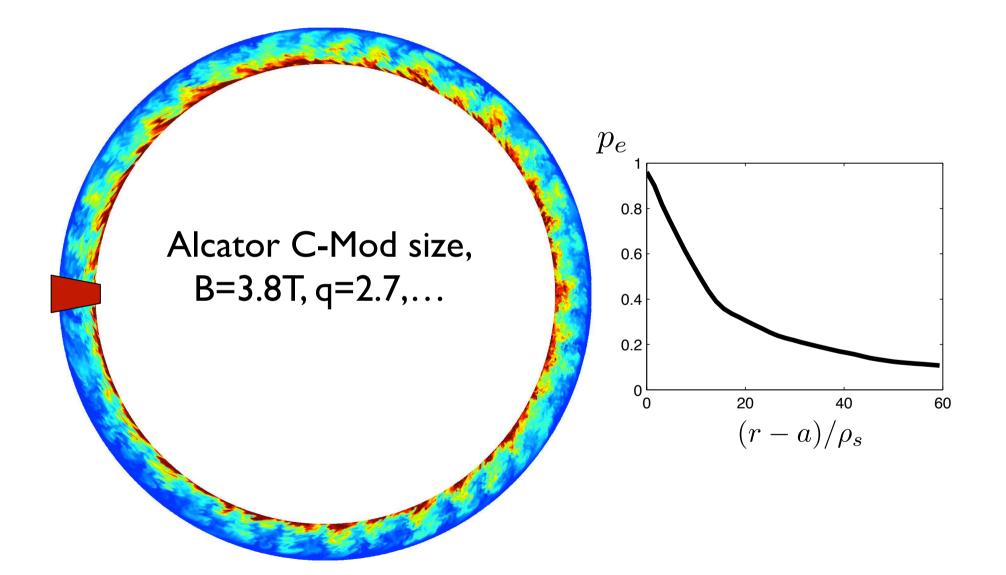
The ballooning scaling, in SI units:

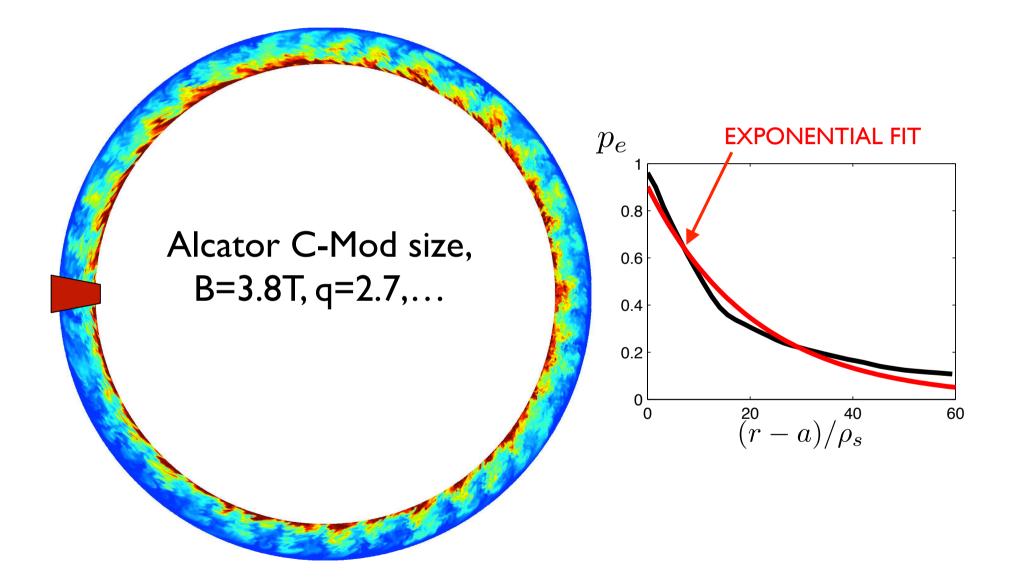
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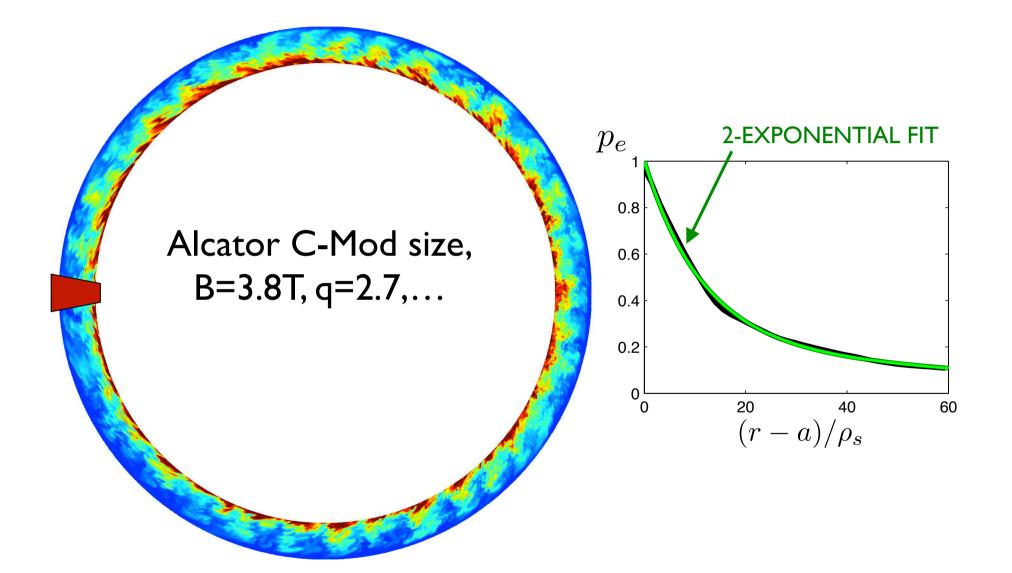


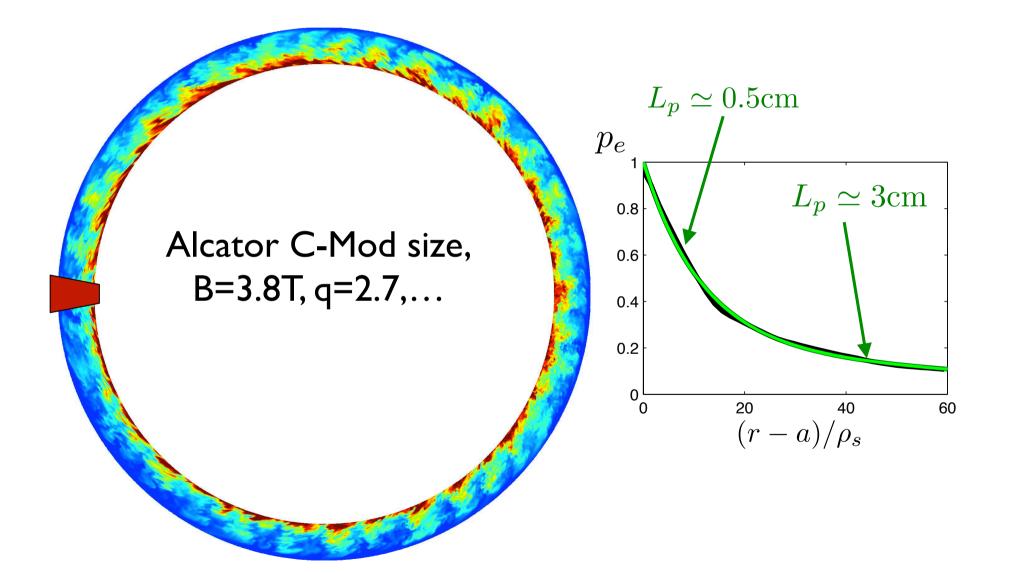
Experimental results provided by G.Arnoux, I. Furno, J.P. Gunn, J. Horacek, M. Kocan, B. LaBombard, B. Labit, and C. Silva.

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How can we approach the SOL width scaling?

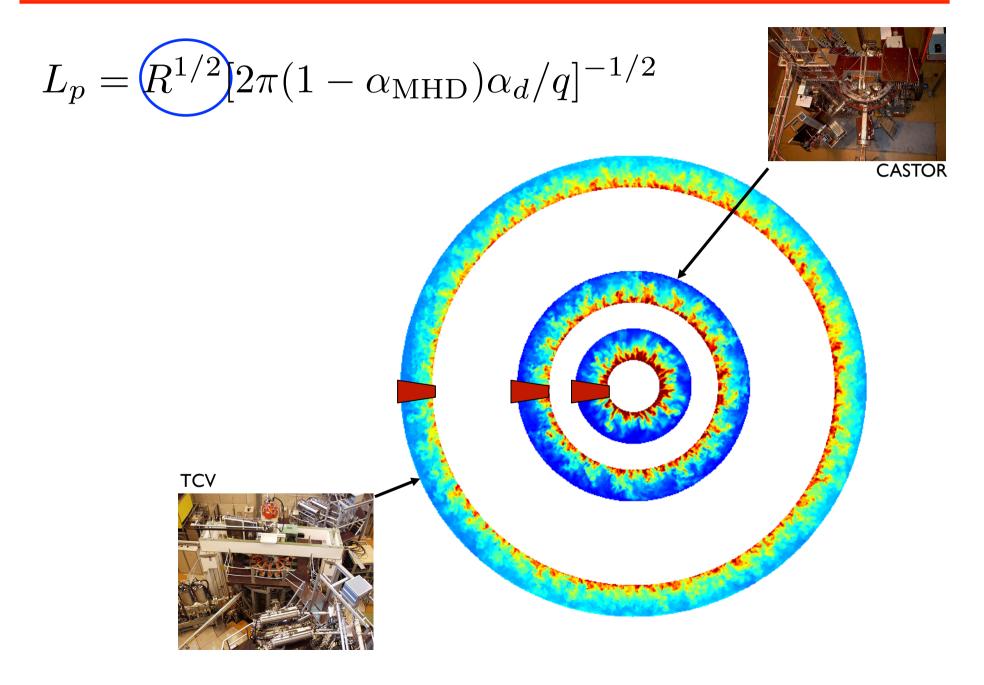
- We can derive a first-principle scaling of the SOL width
- A drift-reduced model is able to represent the main features observed experimentally in the SOL
- Full-size simulations show large fluctuations, intermittent events, large scale ballooning turbulence
- SOL width established from the balance of parallel losses and perpendicular transport, driven by the ballooning instability and saturated by the gradient removal mechanism
- Experimental observations generally in agreement with theoretical observations

http://people.epfl.ch/paolo.ricci

CRPP



Limited SOL width widens with ${\cal R}$



Limited SOL transport increases with β and ν

